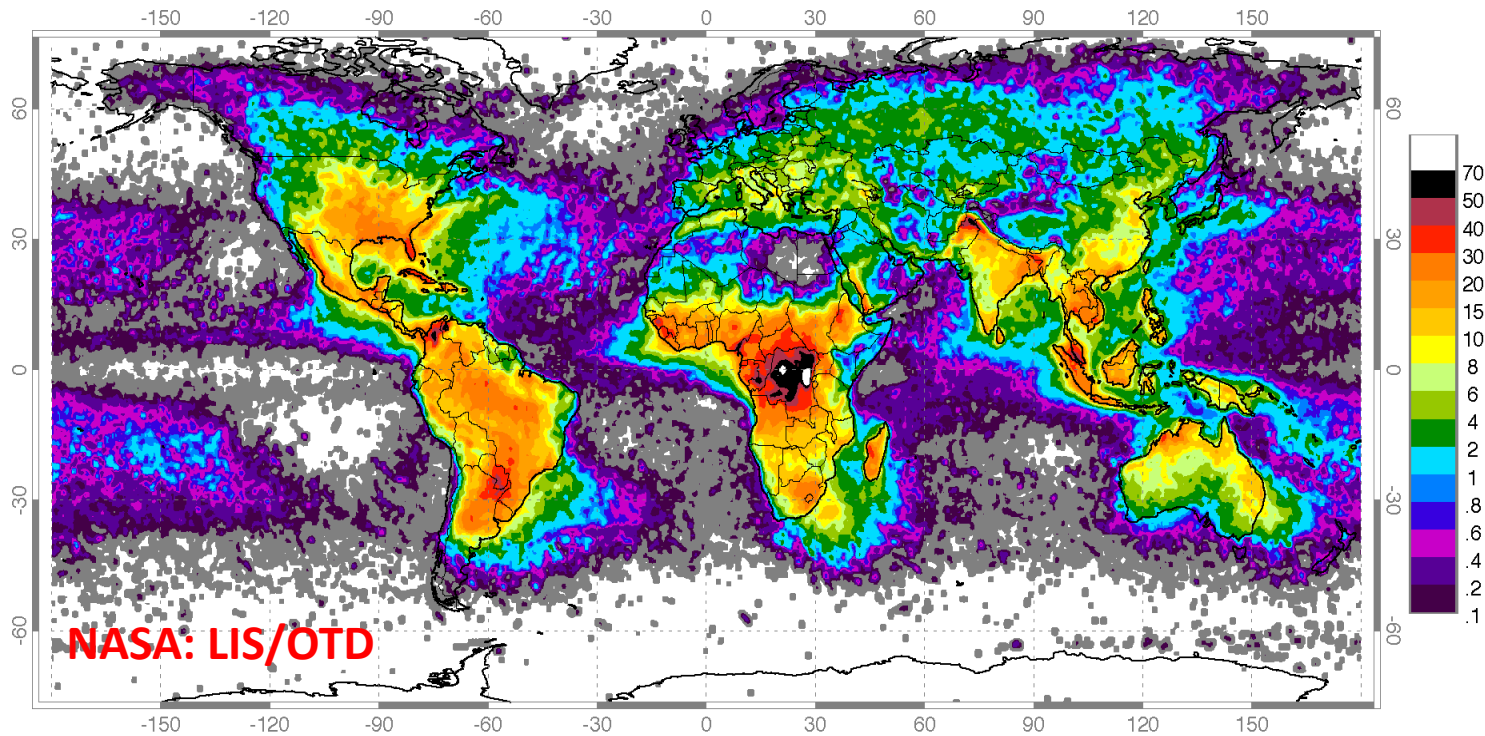




Studying Global Lighting Activity Using Schumann Resonance Measurements



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UJ), Janusz Młynarczyk (AGH) for the Krakow ELF
Research Team



Krajowy Naukowy Ośrodek Wiodący



JAGIELLONIAN UNIVERSITY
IN KRAKÓW

Krakow ELF Research Team

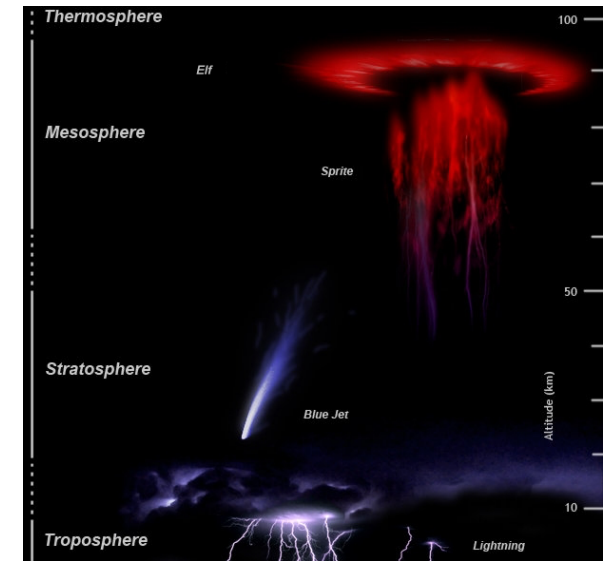
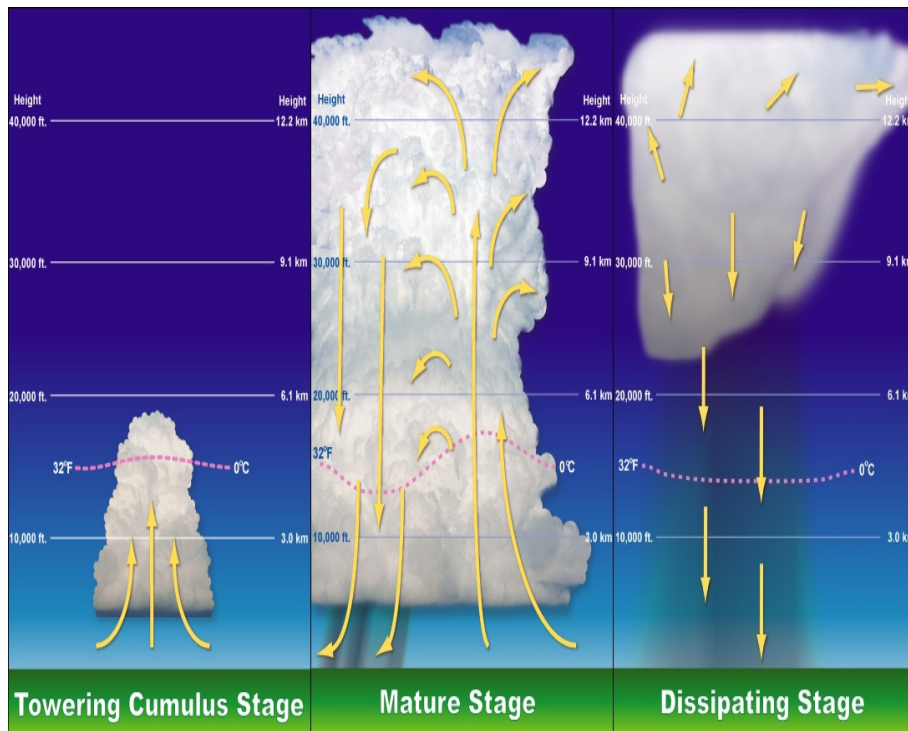
- **Astronomical Observatory UJ:** Andrzej Kułak, Michał Ostrowski, Volodia Marchenko, Jerzy Kubisz (engineer) and Adam Michalec (emeritus)
- **Institute of Physics UJ:** Zenon Nieckarz
- **University of Science and Technology AGH:** Janusz Młynarczyk
- **Institute of Nuclear Physics PAS:** Michał Dyrda
- **International collaborators:**
 - Earle Williams (Massachusetts Institute of Technology)
 - Gabriella Satori (Hungarian Academy of Sciences)

Storms in Nature

Typically – 1800 thunderstorms all over the World and 50-100/sek lightning discharges

Conditions to form thunderstorm:

1. Moisture
2. An unstable air mass
3. A lifting force (heat)



Lightning discharges:

- Cloud-to-ground (CG)
90 of all discharges CG⁻
typical current $I_{CG^-} = 22 \text{ kA}$ but $I_{CG^+} = 300 \text{ kA}$
- Intra-cloud (IC)
- Cloud-to-cloud (CC)

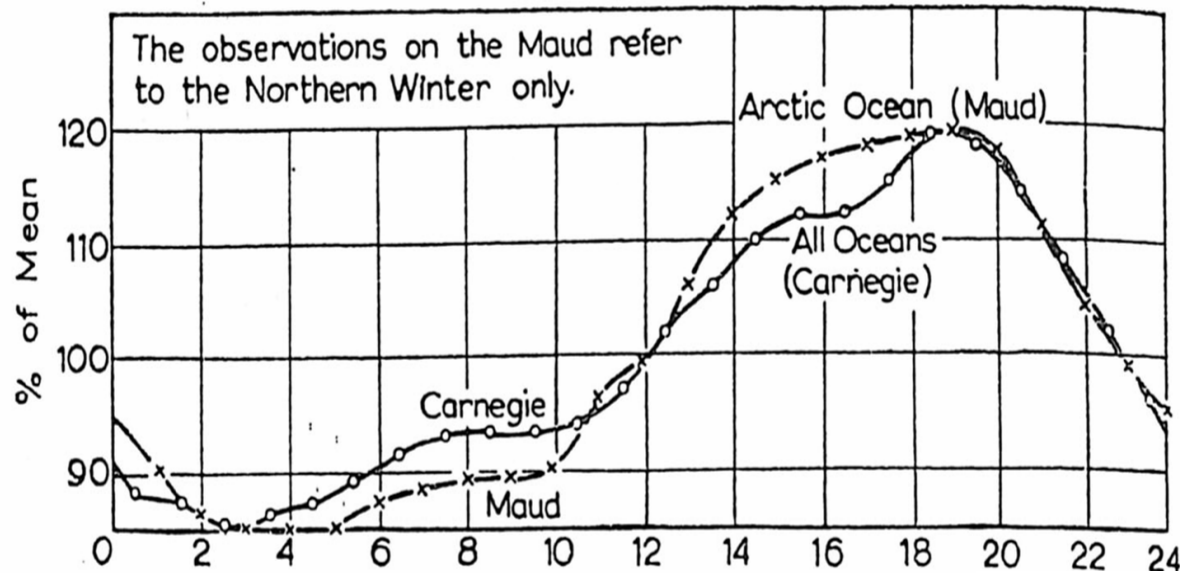
Also transient luminous event (TLE) :

- Sprites
- Jets
- Elves
- Terrestrial Gamma Ray flashes (TGF)



Carnegie Curve

1752 Lemonnier: in fair weather regions there is a persistent E-field of 100 V/m pointing downwards



1920s: diurnal variation of the atmospheric electric field, independent of location and local time, but dependent only on **universal time** !

Global Electric Circuit:

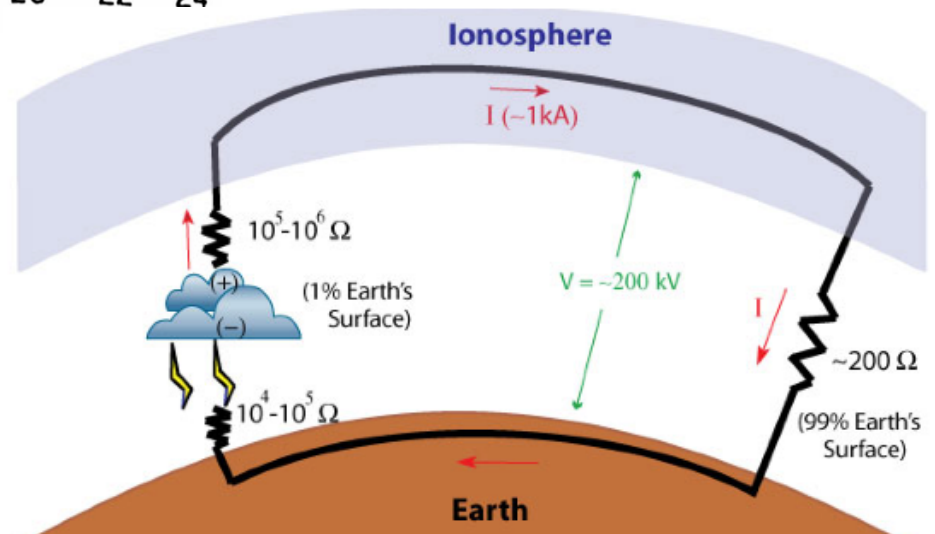
Earth's charge $Q = 4\pi R^2 \epsilon_0 E \approx 4.5 \times 10^5$ Coulomb

capacitance of the Earth $C = Q/V \approx 1$ Farad

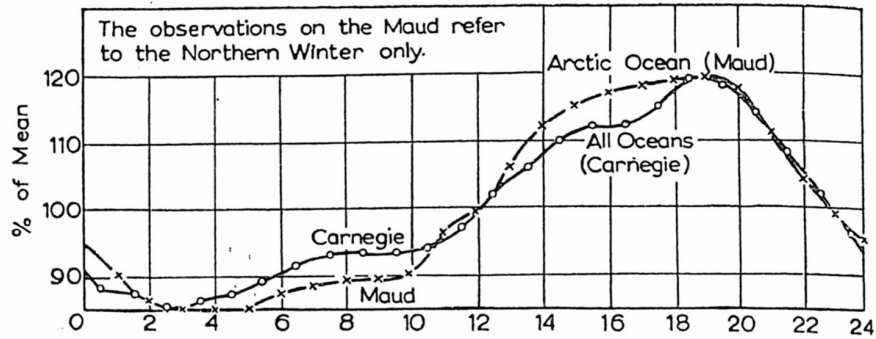
time scale for the discharging of the capacitor:

$\tau = R \cdot C = 300$ sec!

Where are the batteries !?



Where are the batteries ?



Batteries = tropical thunderstorms

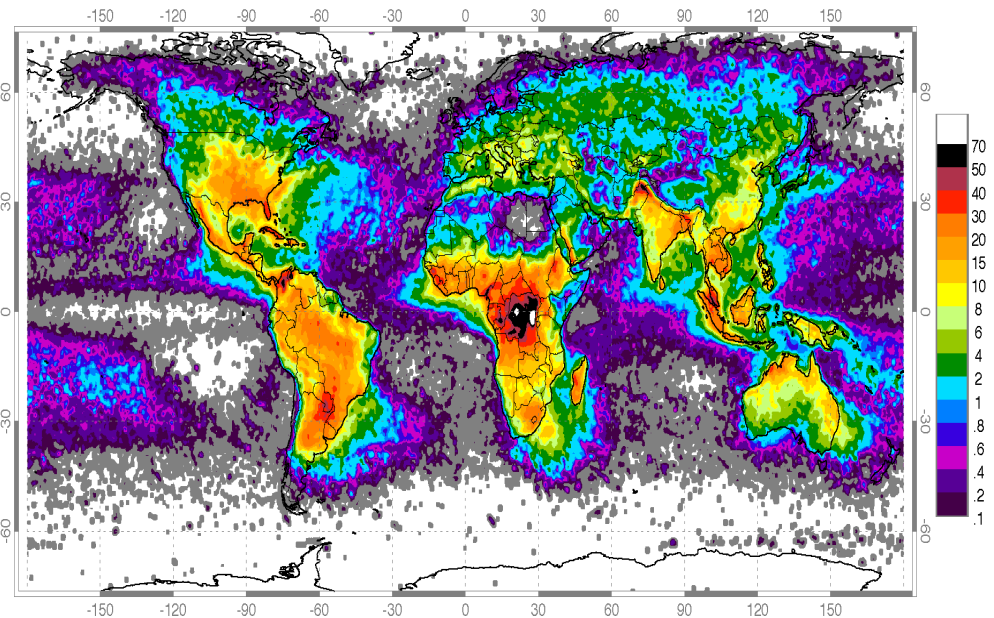
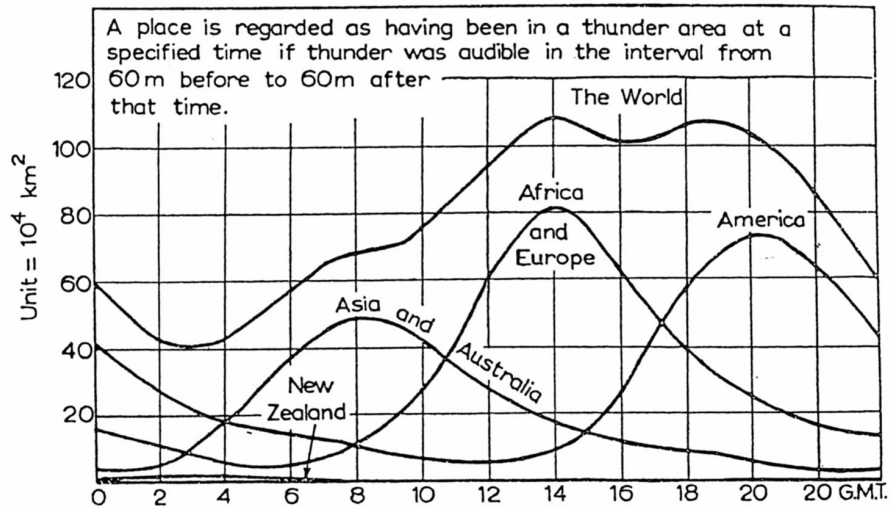


FIG. 54. Diurnal variation of land thunder areas and of potential gradient in unperturbed conditions. (From WHIPPLE and SECRASE, 1936, Figs. 9 A and B, p. 19.)

Lightning detection methods

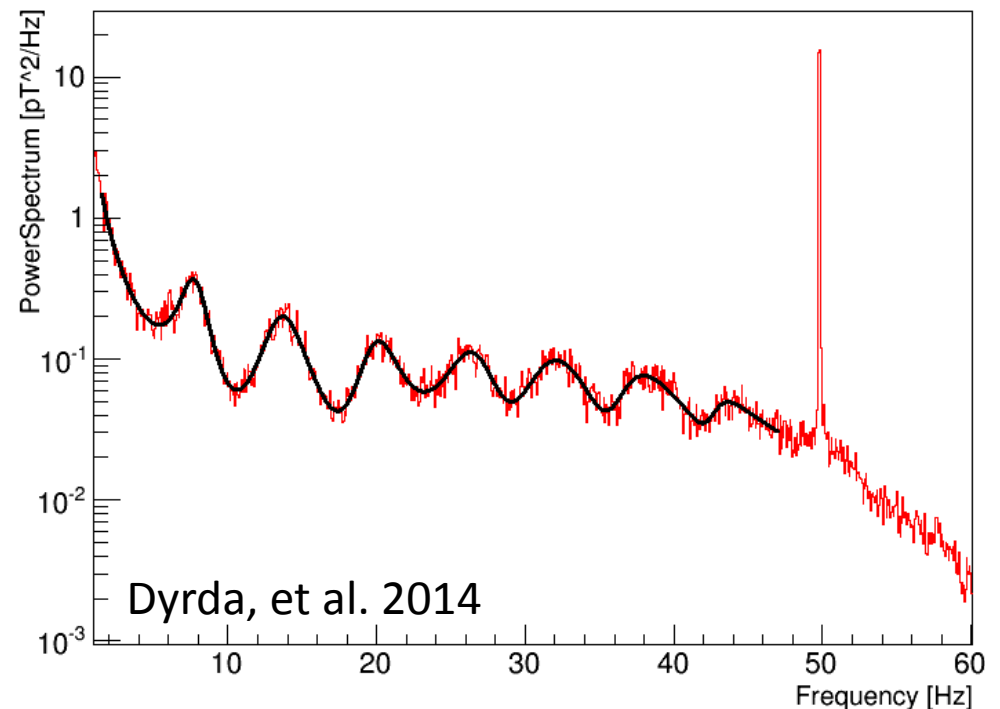
- Optical observation done by satellites:
 - detailed information about observed area (10^5 km²), (i.e. LIS/OTD NASA)
 - new system of three satellites in 2018
 - but clouds screening – might not be able to measure total storms activity
- Radio observations: VLF (3-30 kHz), HF and VHF:
 - detection range varies from ~100 km up to ~1000 km (VLF) (global – WWLLN; local – i.e. LINET in Europe)
- ELF observation (1Hz-3kHz):
 - single detector can detect signals from all atmospheric lightning's
 - unambiguous map construction of global storms activity up to distances ~ 20 Mm!
 - with few detectors -> possibility to image global storm activity in geographical coordinates
 - can measure CG lightning's, no screening effect

Extremely Low Frequency Waves - Schumann Resonance

- Global electromagnetic resonance – solutions to the Maxwell equations - Winfried Otto Schumann (1952)
- Earth (limited dimensions) → resonance cavity → excited by lightning → antennas that radiate electromagnetic energy
- Signals very weak at large distances from sources
- Earth-ionosphere waveguide – resonator at ELF frequencies
- Amplifies spectra signals at resonance frequencies (ideal cavity - no damping):

$$f_n = \frac{c}{2\pi a} \sqrt{n(n+1)} = 10.6, 19.3, \dots [\text{Hz}]$$

(Real) Earth-ionosphere cavity:
 $f_n = 8, 14, 20, 26 \dots \text{Hz}$

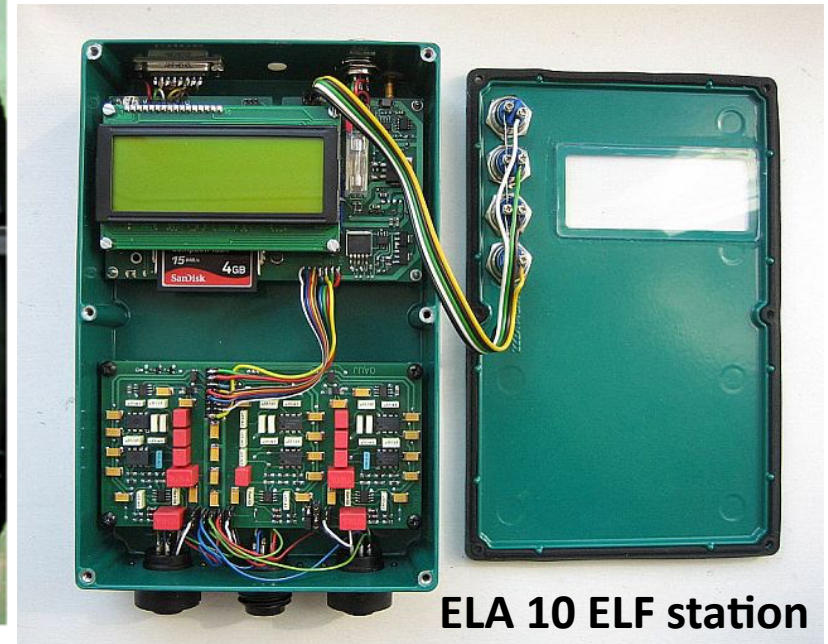
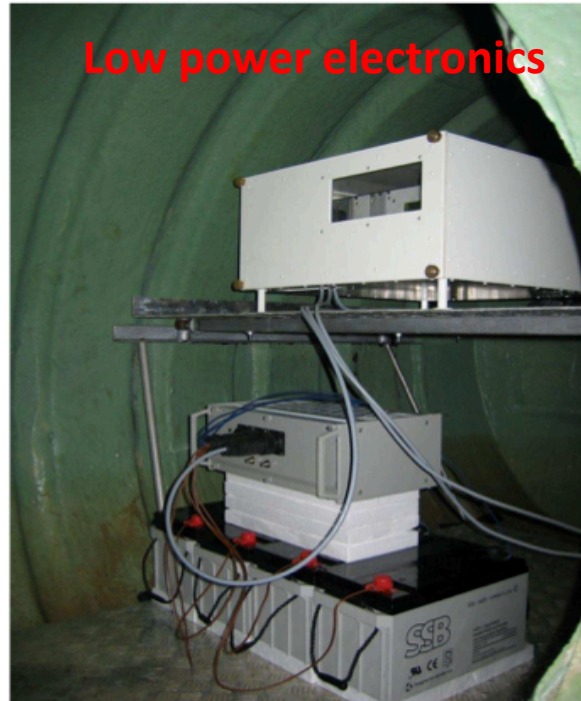
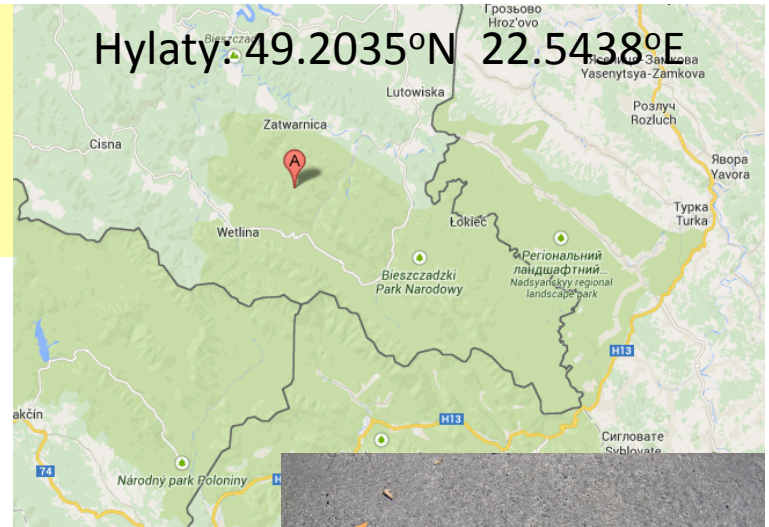


ELF electromagnetic field measurements at the Hylaty station in Poland

Hylaty ELF station specifications [Kulak et al., 2014]:

- Two ELF receivers: 0.03 - 55 Hz and 0.03 - 300 Hz
- Two signal channels NS and EW for each receiver
- Sampling frequencies: 175 and 900 Hz
- Sensitivity: 0.05 pT/sqrt(Hz) at 10 Hz
- Data storage – CF card 32Gb

Also special GPS time units



ELA 10 ELF station

Spectral Decomposition Method (Decomp)

Kulak et al. 2006

$$W(f, \theta) = s + \frac{z}{f^m} + \sum_{k=1}^N \frac{p_k(\theta)[1 + e_k(\theta)(f - f_k)]}{(f - f_k)^2 + g_k^2/4}$$

$W(f, \theta)$ – wave power spectrum [T^2/Hz]
 s – white noise term [T^2/Hz]
 z/f^m – color noise term, $z \rightarrow [\text{Hz}^{m-1}T^2]$
 p_k – power parameter of k^{th} mode [$T^2\text{Hz}$]
 e_k – asymmetry parameter [Hz^{-1}]
 f_k – resonant (intrinsic) frequency [Hz]
 g_k – peak width [Hz]
 $N = 7$ in our case

Lack of asymmetry
 $e_k = 0$

Lorentzian Curve
but $f_k \rightarrow f_k^*(\theta)$
all parameters – classical meanings

Parameter m :
 $\langle m \rangle = 2.06$

Thunderstorms activity maps - algorithm

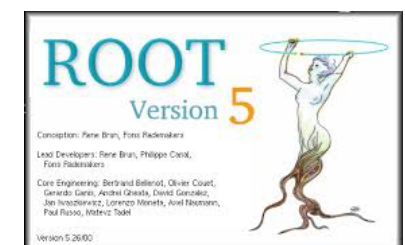
- Time bins – 10 min
- FFTW – observational SR power spectrum
- Decomp – in frequency range 1-48 Hz



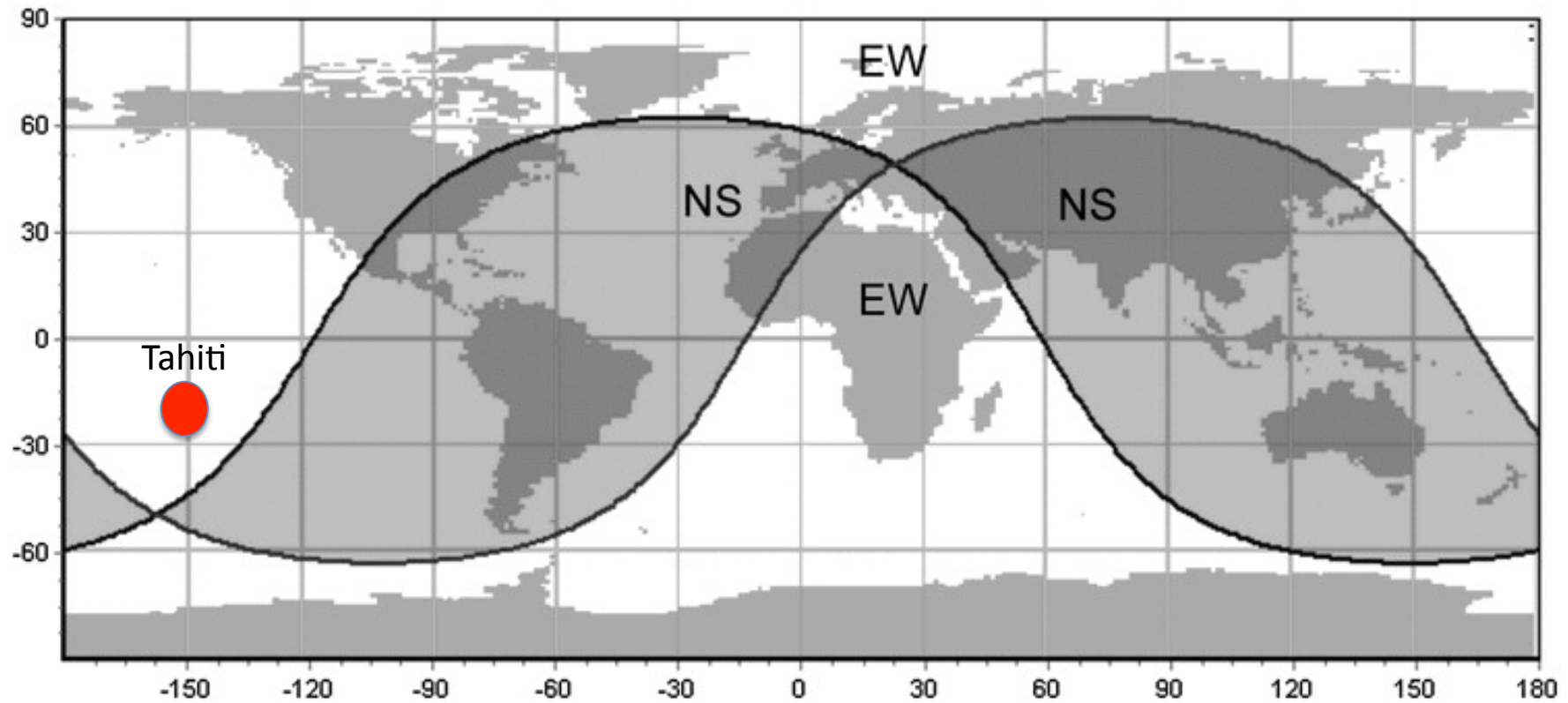
FFTW

$$W(f, \theta) = s + \frac{z}{f^m} + \sum_{k=1}^N \frac{a_k [1 + e_k (f - f_k)]}{(f - f_k)^2 + (\frac{g_k}{2})^2}$$

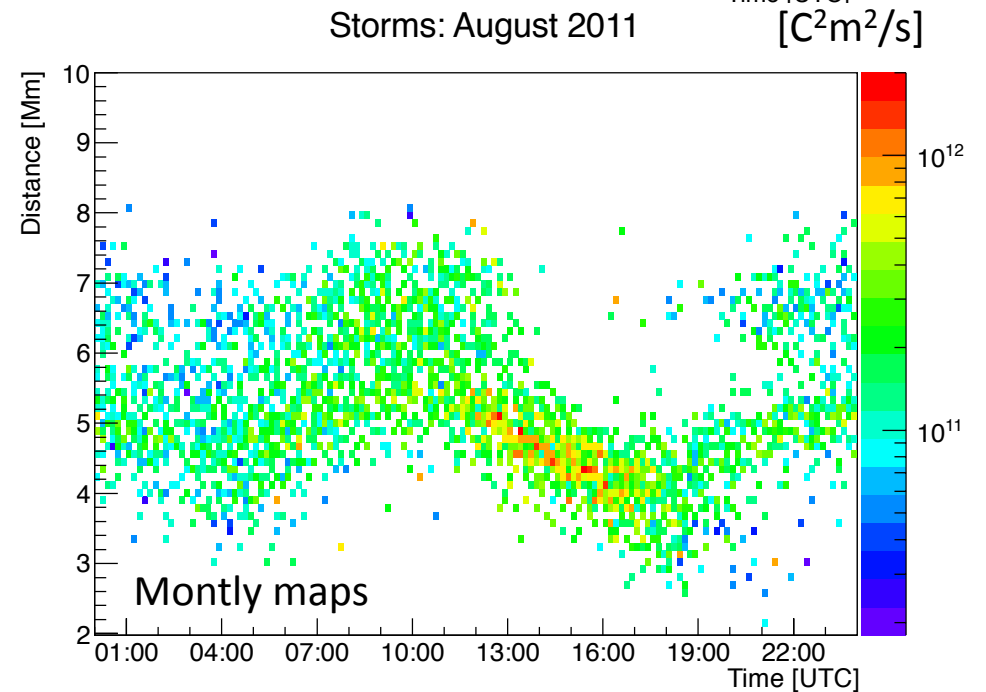
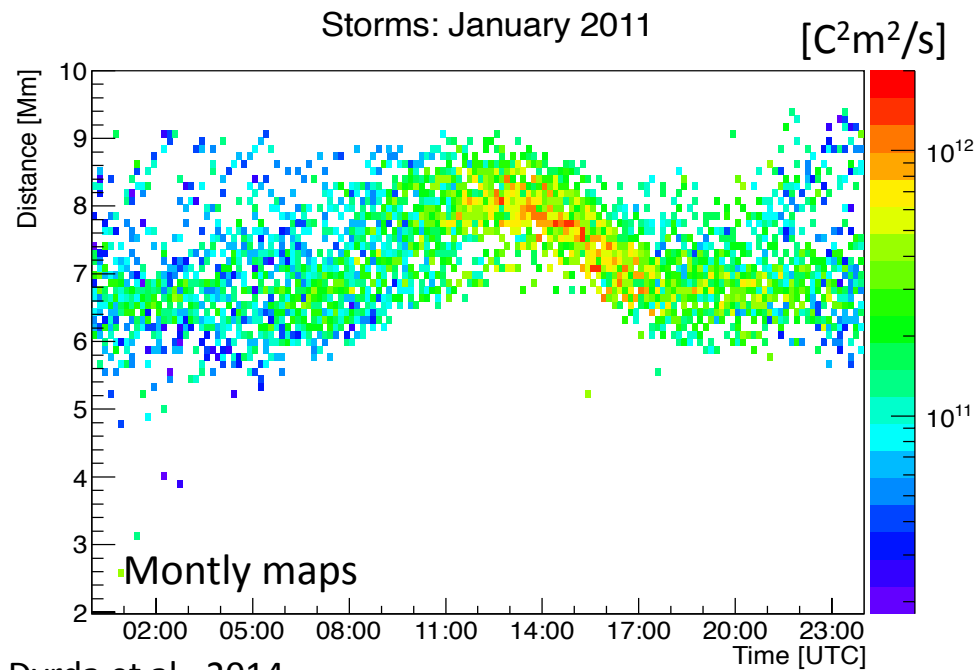
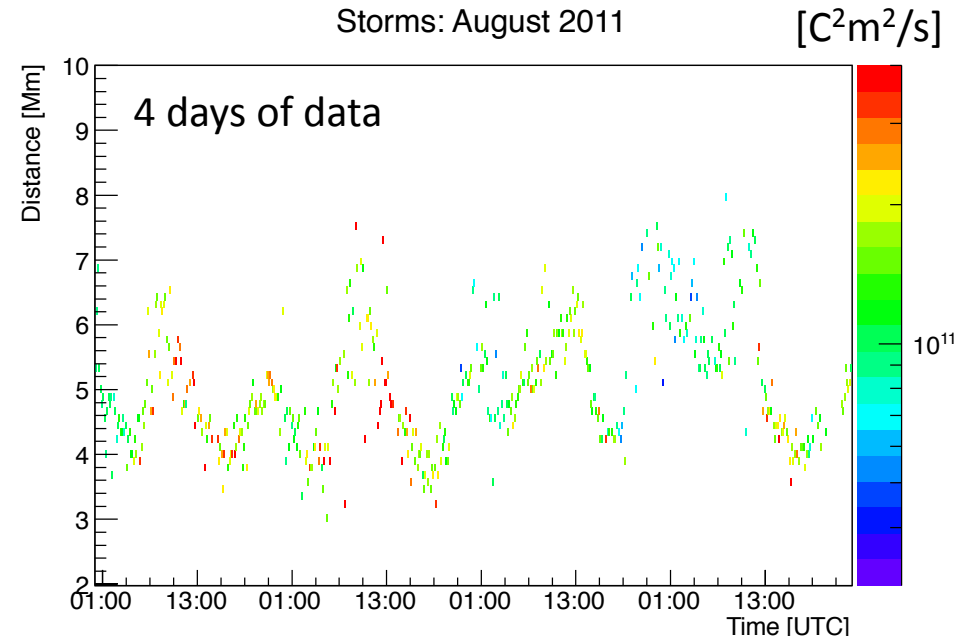
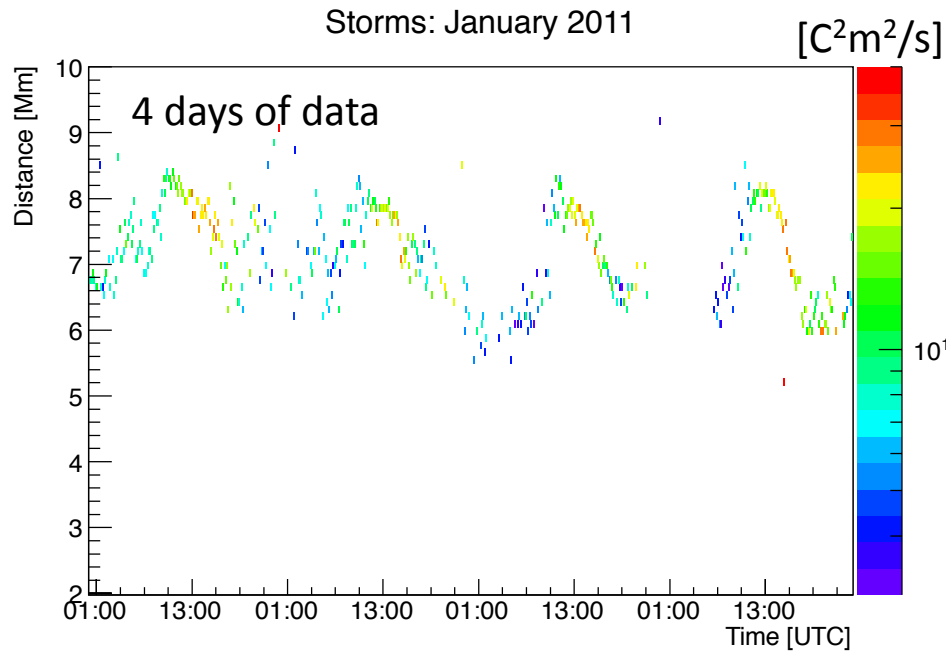
- Model power spectra calculated for known source-observer separations with step ~ 0.1 Mm
- χ^2 test \rightarrow distance to the thunderstorm center [Mm]
- Thunderstorm center intensity in absolute units $(C \cdot m)^2/s$
- Data storage – Cyfronet
- Thunderstorms activity maps calculated using PLGrid infrastructure
- Our codes (C++) and ROOT – data analysis framework



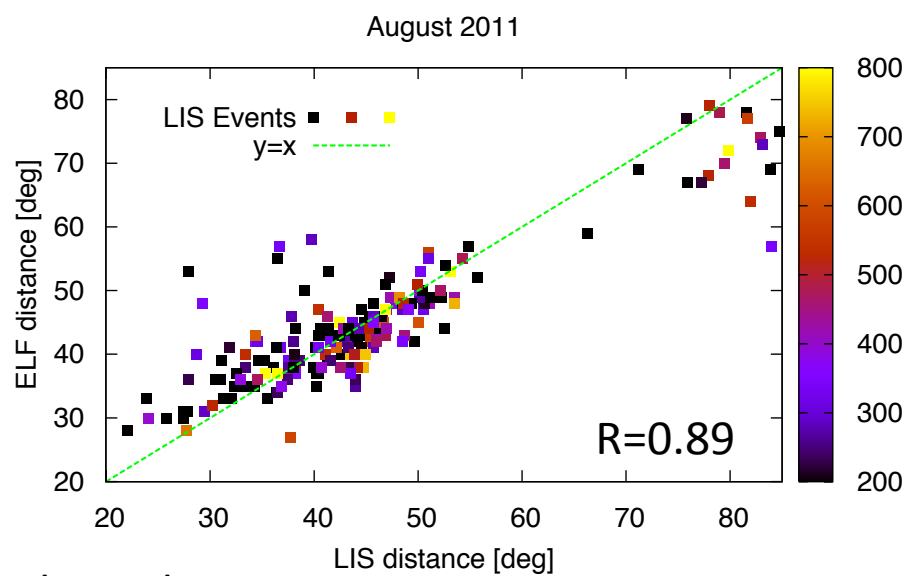
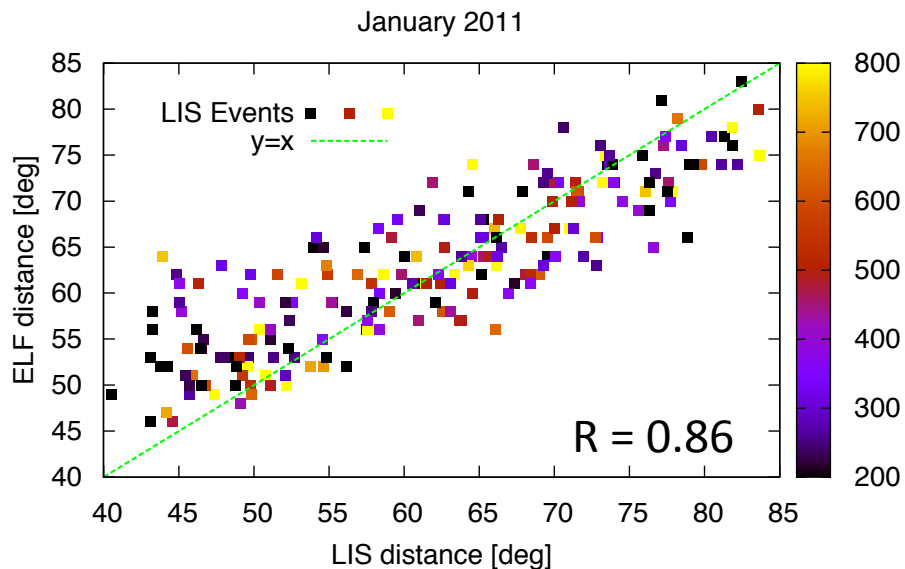
Hylaty – great circles



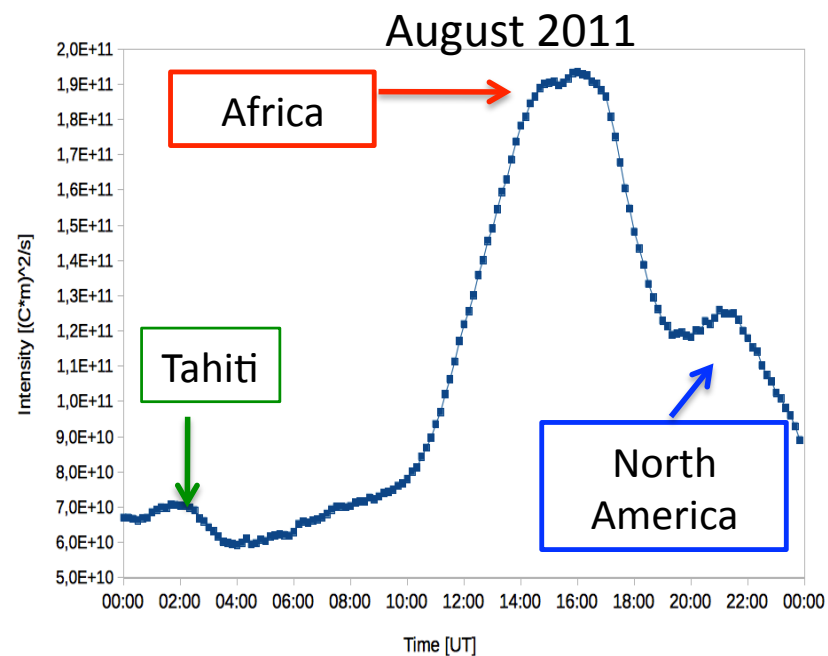
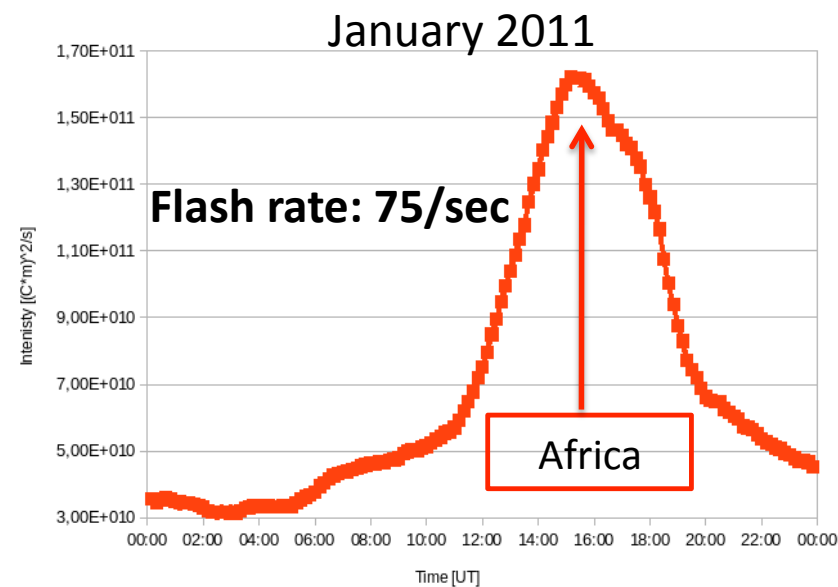
African thunderstorms activity in 2011



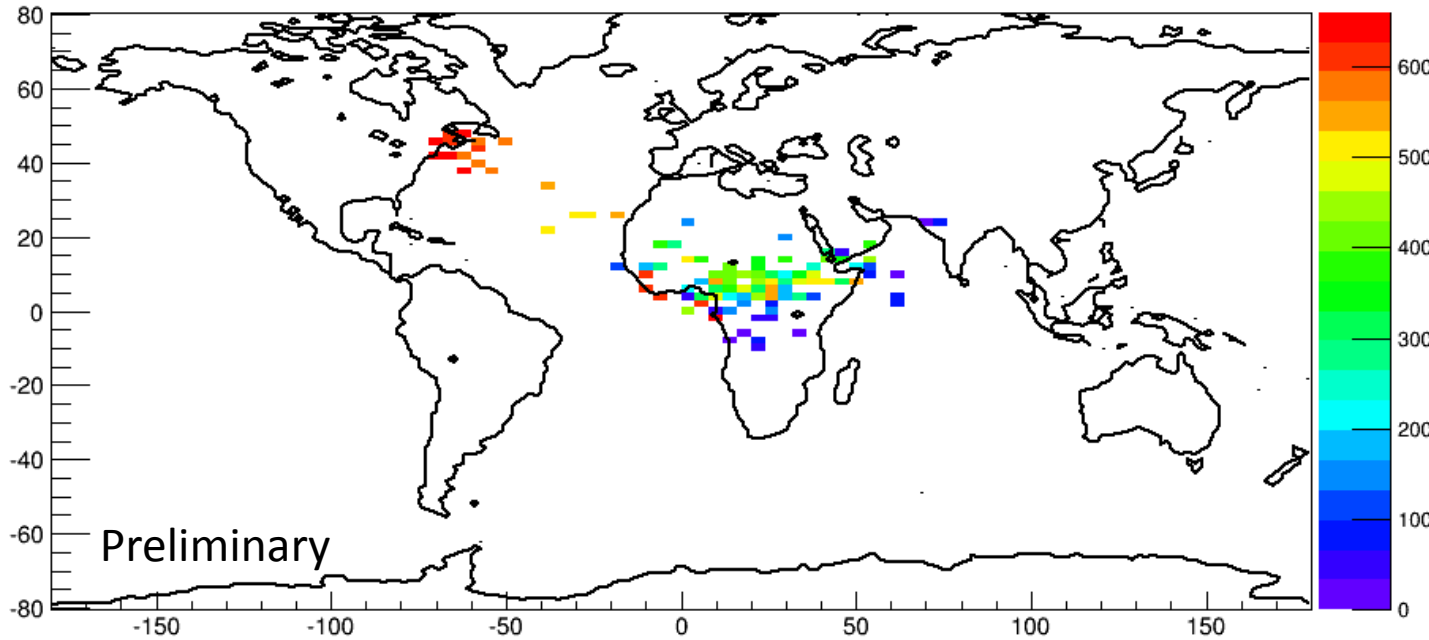
Correlation with optical measurements



Our carnegie-type curves

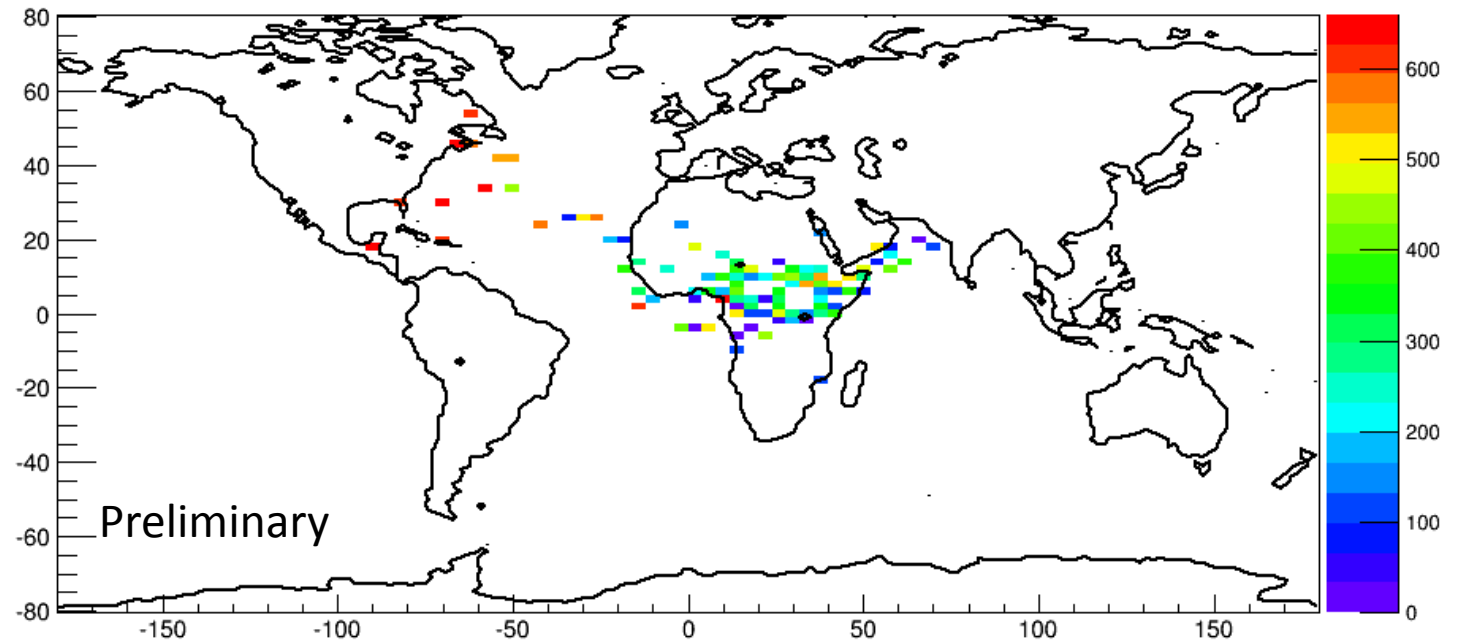


Hylaty 300 Hz: 2013.07.31: 10-21 UT



Thunderstorm
centers in
geographical
coordinates

Wigry 300 Hz: 2013.07.31: 10-21 UT



In summer 2013
simultaneous
observation with 3
ELF station:

- Hylaty
- Wigry Lake
- Karkonosze

ELF Very Long Baseline Radio Locator



Backup slides

Schumann resonance – solar flares

